How the Earth's Rotation Affects the Weather? by **Jun-Ichi Yano Meteo France** Toulouse

Original "Zapytaj Fizyka" Question:

What happens to the weather when the Earth begins to rotate slower/faster?



Question of Taxonomy:

cf., M. Foucaullt: *Les Mots et les Choses*





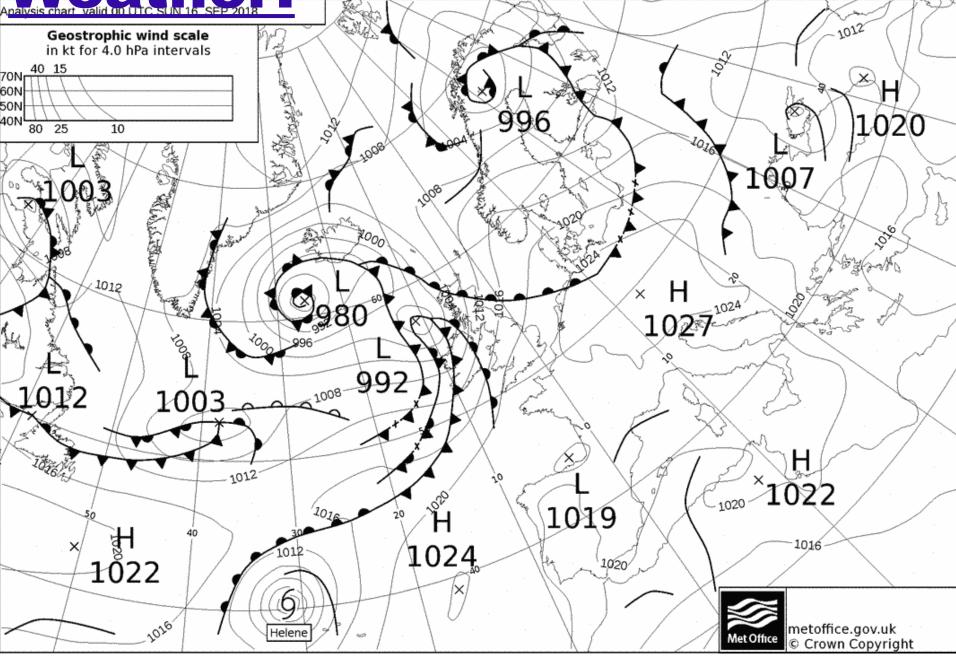


Weather?

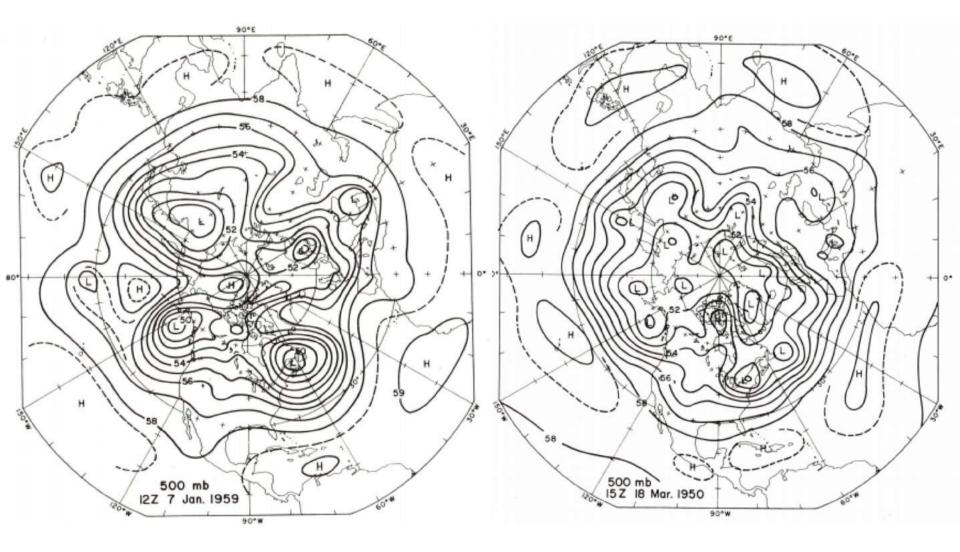




Archived by www.wetter3.de

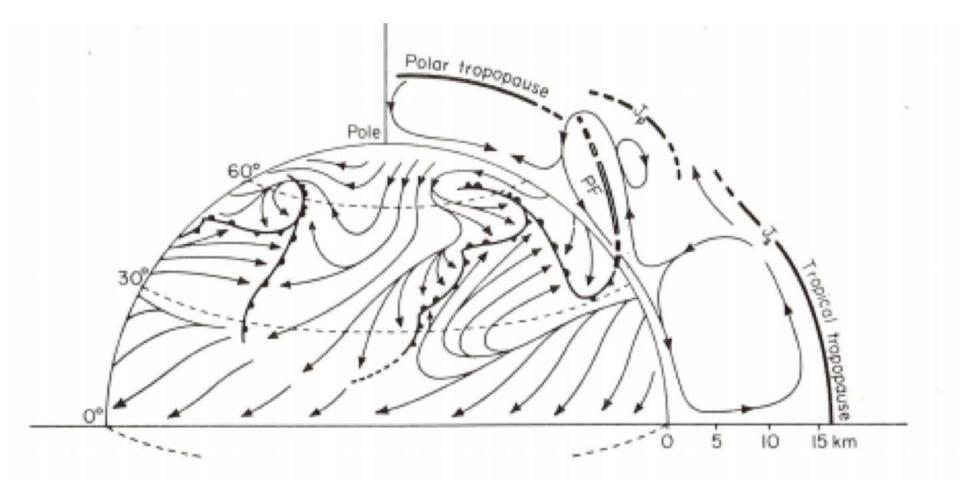


Weather (Global View):



(Palmén and Newton 1969)

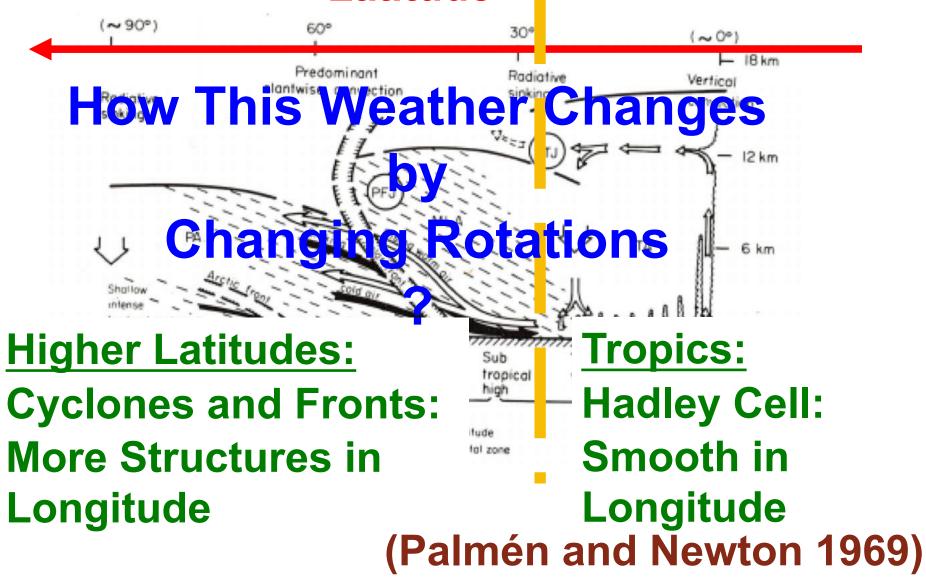
Weather (Global View):



(Palmén and Newton 1969)

Weather(Global View):





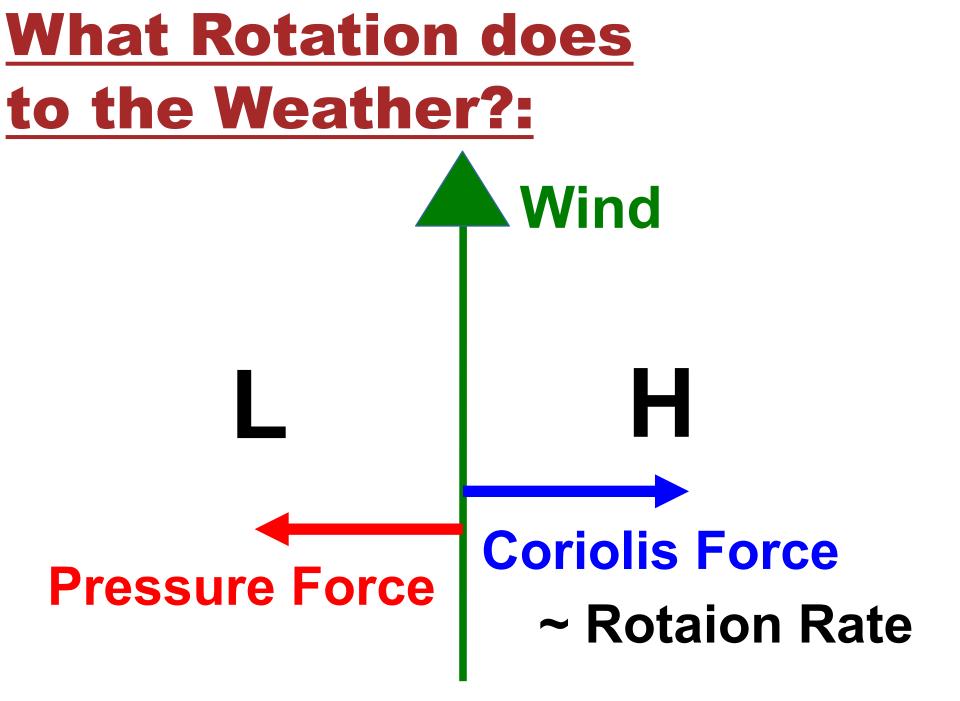
What Rotation does to the Weather?:

Н

Winds~Pressure Lines:

Coriolis Force

Н

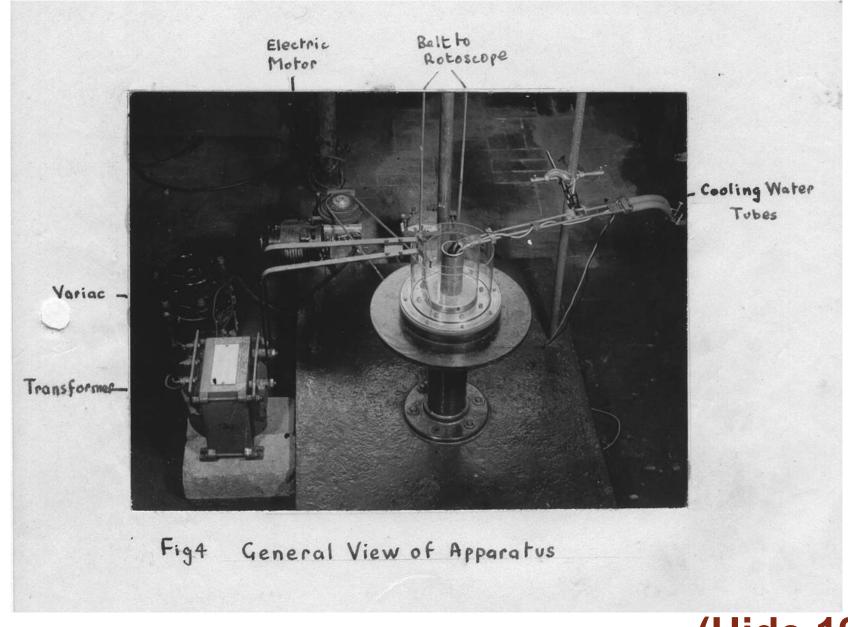


Q: What happens to the weather when the Earth begins to rotate slower/faster?

Three Approaches:

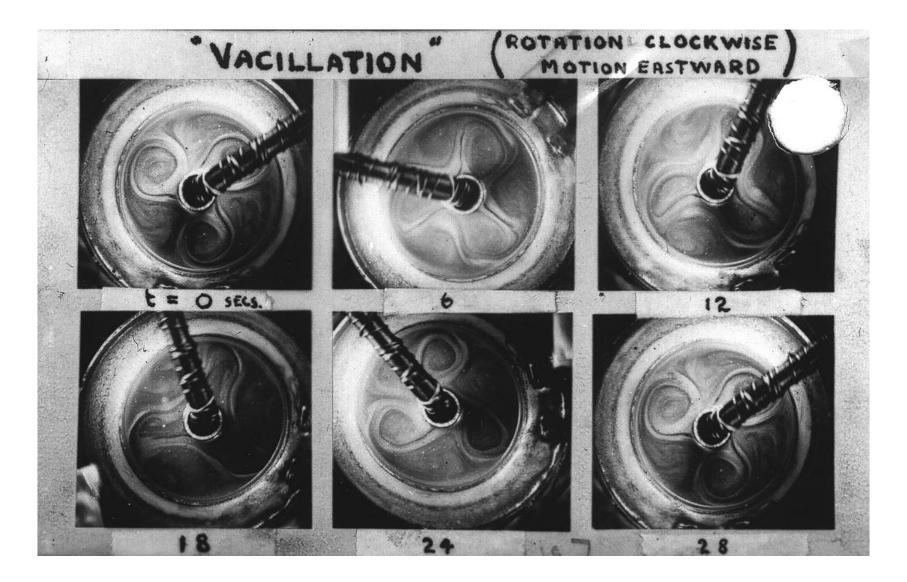
- **1. Laboratory Experiments**
- 2. Planetary Atmospheres
- 3. Theory: Explanations

Laboratory Experiment: Rotating Tank:



(Hide 1958)

Laboratory Experiment: Snap Shots:

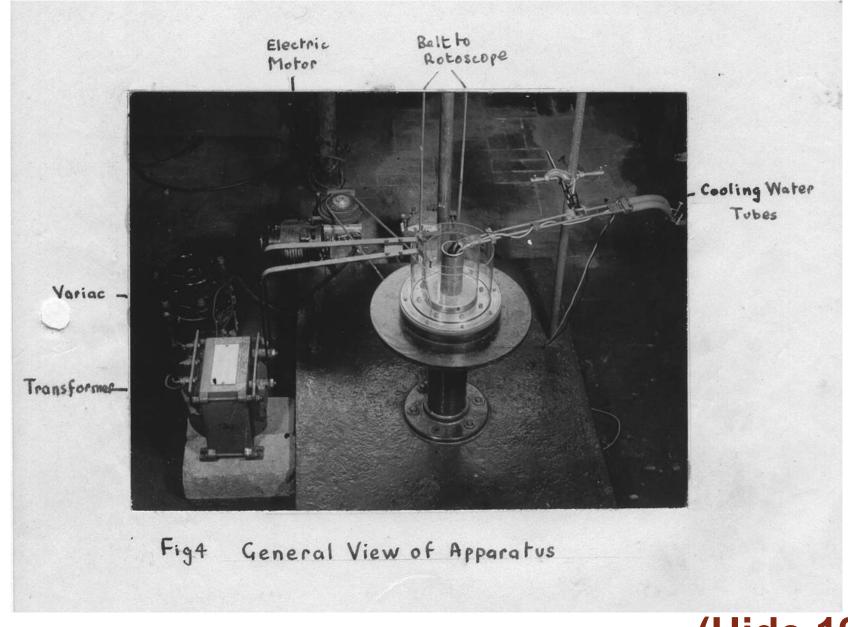


(Hide 1958)



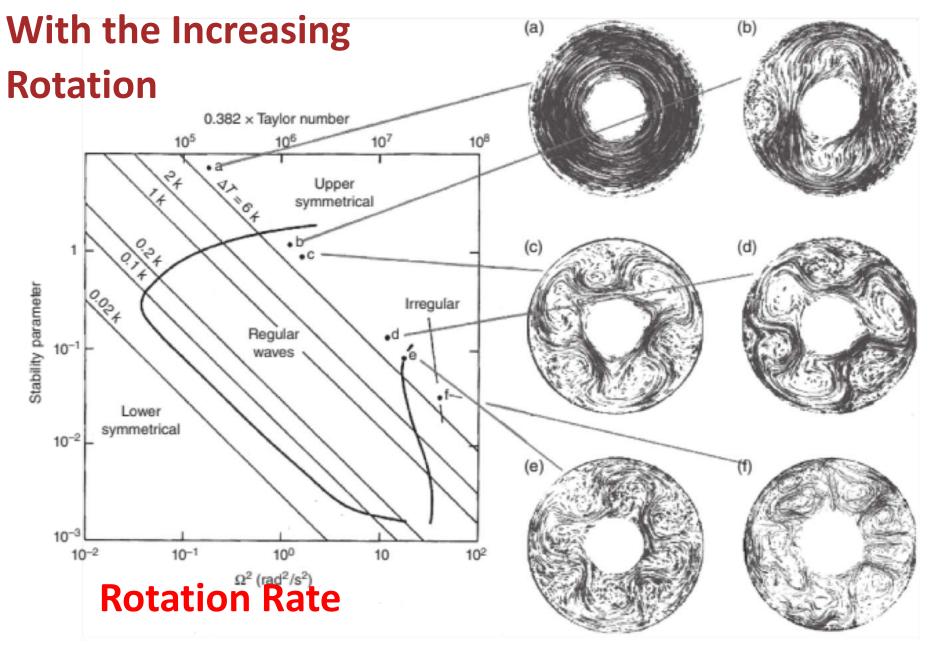


Laboratory Experiment: Rotating Tank:



(Hide 1958)

Rotating Tank Experiment:



Rotating Tank Experiment:

- **Summary of the Results:**
- With Very Slow Rotations: Smooth in Longitudinal direction: Hadley Cell

- With Increasing Rotations:
- **Breaks down into**
- **Eddies in Longitudinal direction:**
- Weather of Cyclones and Fronts
- With Further Increase of Rotations:
- Weather Pattern becomes increasingly Irregular

Planets with Atmospheres



Venus



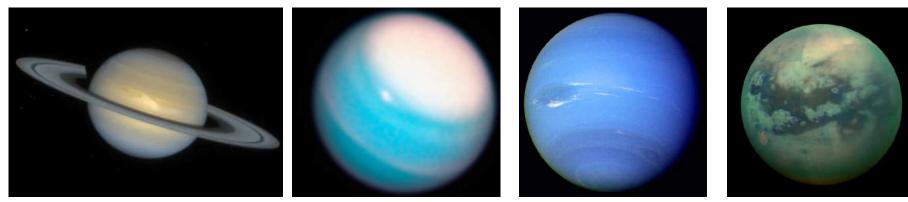
Earth



Mars



Jupiter

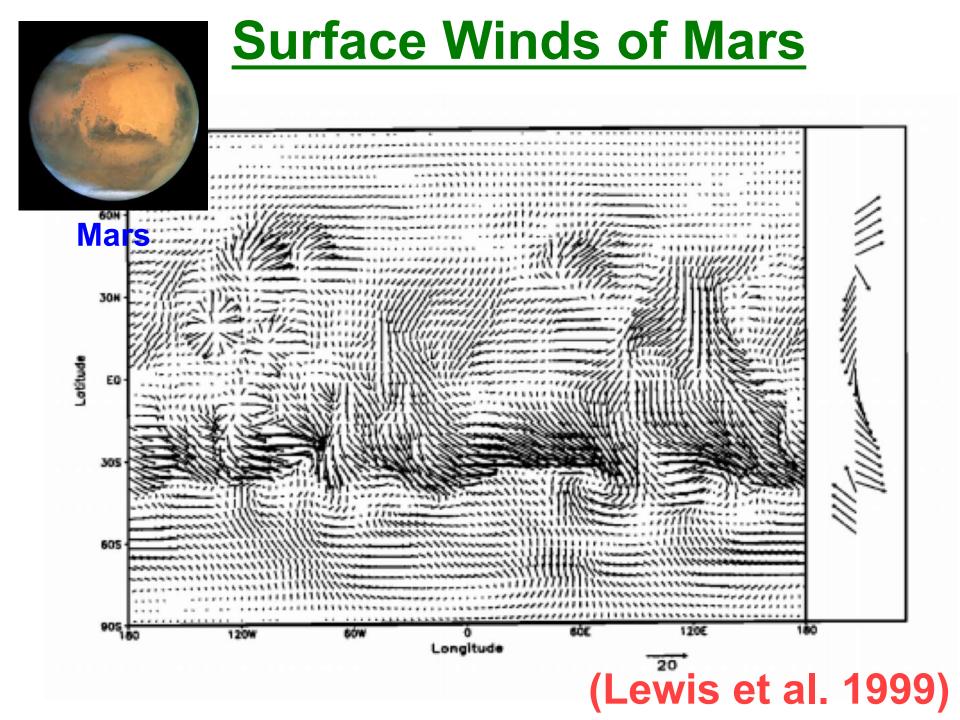


Saturn

Uranus

Neptune

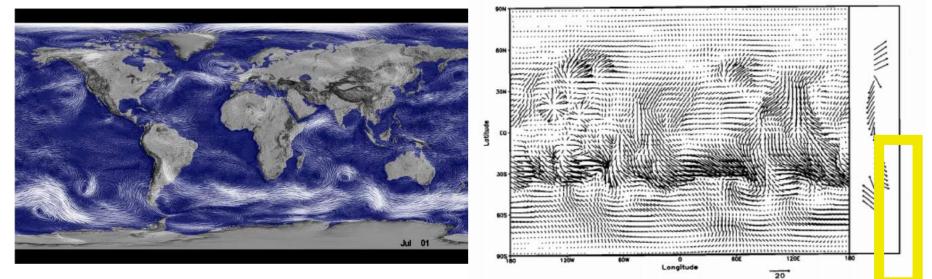
Titan (Saturn's Satellite)



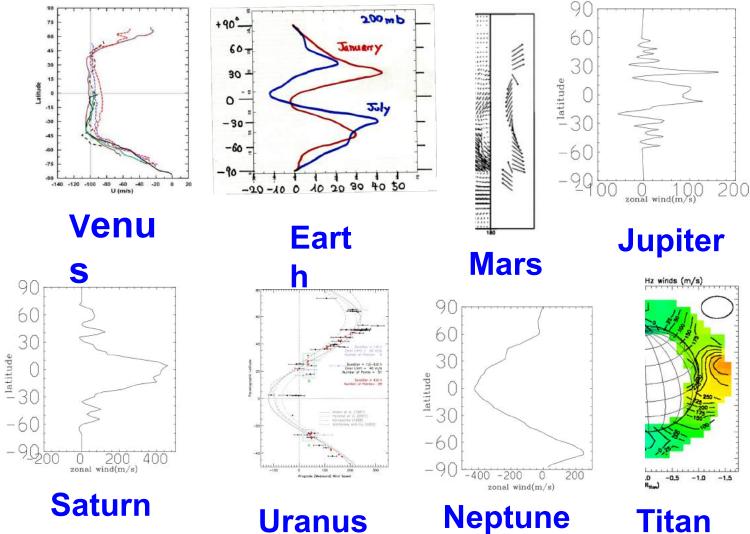
Surface Winds: Earth and Mars







Eastward Winds with Latitudes:

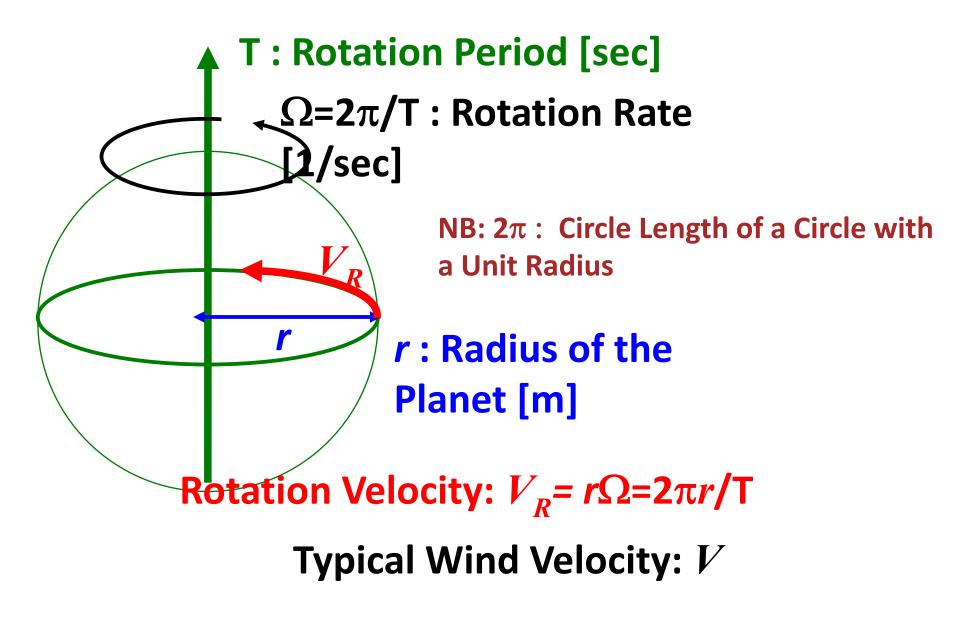


(Saturn's Satellite)

Basic Characteristics of the Planets:

Planet	Rotation Period (re	Radius elative to Earth)	Typical Wind Speed
Venus	-243 days	0.95	100m/s
Earth	24 hours	1.0 (6371 km)	10m/s
Mars	25 hours	0.53	10m/s
Jupiter	10 hours	11	100m/s
Saturn	10 hours	9.4	200m/s
Uranus	18 hours	4.1	100m/s
Neptune	18 hours	3.9	200m/s
Titan	383 hours	0.4	200m/s

Degree of Importance of Rotation?



T : Rotation Period [sec]

 $\Omega = 2\pi/T$: Rotation Rate [m/sec]

r : Radius of the Planet [m]

Rotation Velocity: $V_R = r\Omega = 2\pi r/T$

Typical Wind Velocity: V

Degree of Importance of Rotation?:

Typical Wind Velocity/Rotation Velocity= $V/V_R = V/r$ Ω : Rossby Number

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Importance of Rotation?:Rossby number

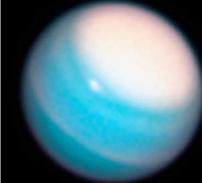
Planet	Rotation Period	Rossby number	
Venus	-243 days	-55	
Earth	24 hours	0.02	
Mars	25 hours	0.04	
Jupiter	10 hours	0.008	
Saturn	10 hours	0.02	
Uranus	18 hours	0.04	
Neptune	18 hours	0.08	
Titan	383 hours	17	

Planets with Atmospheres In Increasing Order of the Rossby number:









Jupiter



Earth

Uranus





Mars

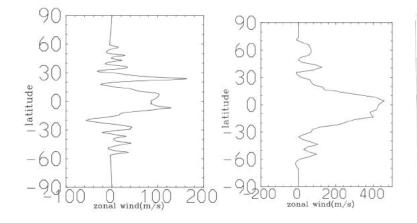


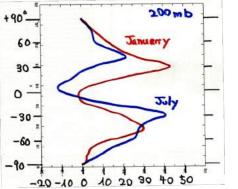
Titan (Saturn's Satellite)

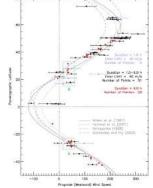


Venus

Eastward Winds with Latitudes: In Increasing Order of the Rossby number:









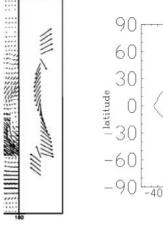


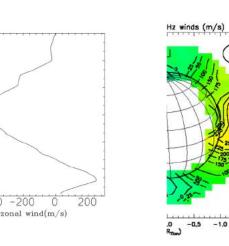
-200

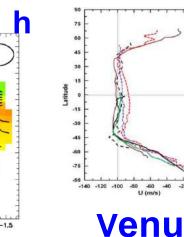
Neptune











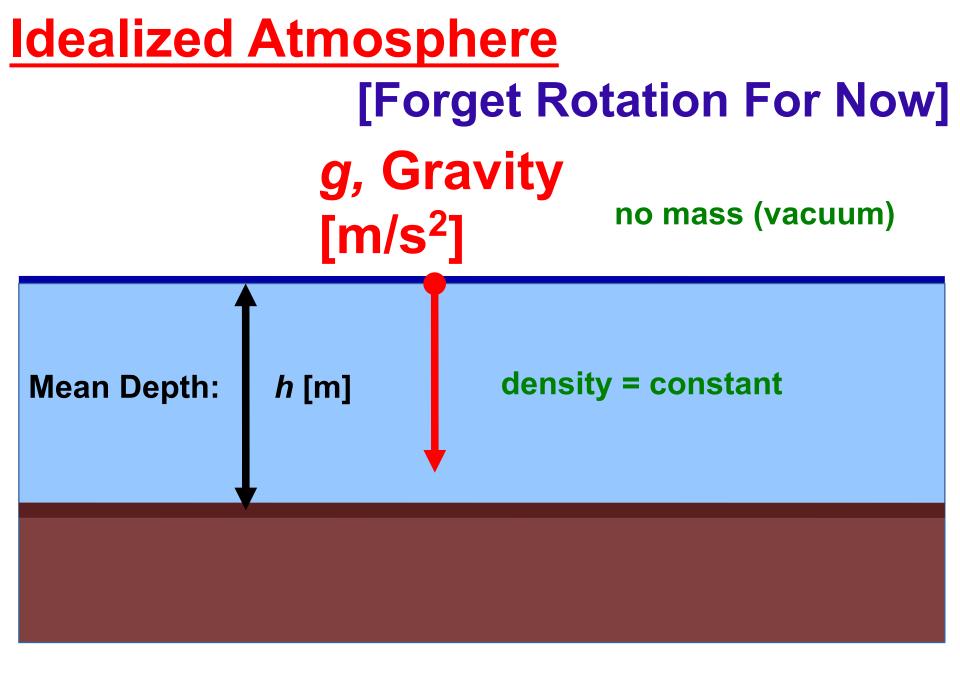
-20



Titan S (Saturn's Satellite)

- Eastward-Wind Profile becomes Finer when the atmosphere is More strongly controlled by Rotation
- Eastward-Wind Profile becomes Broader when the atmosphere is Less strongly controlled by Rotation

But Why?

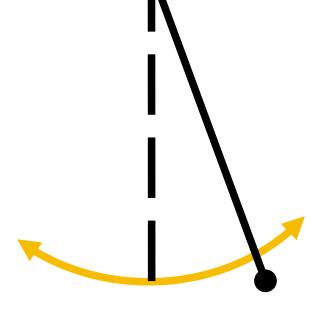


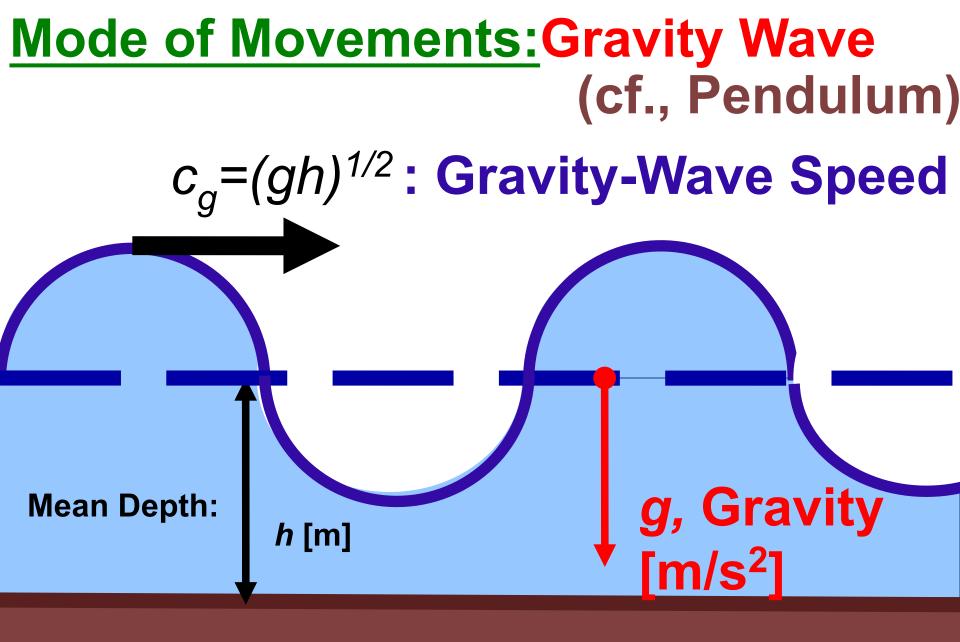
Idealized Atmosphere

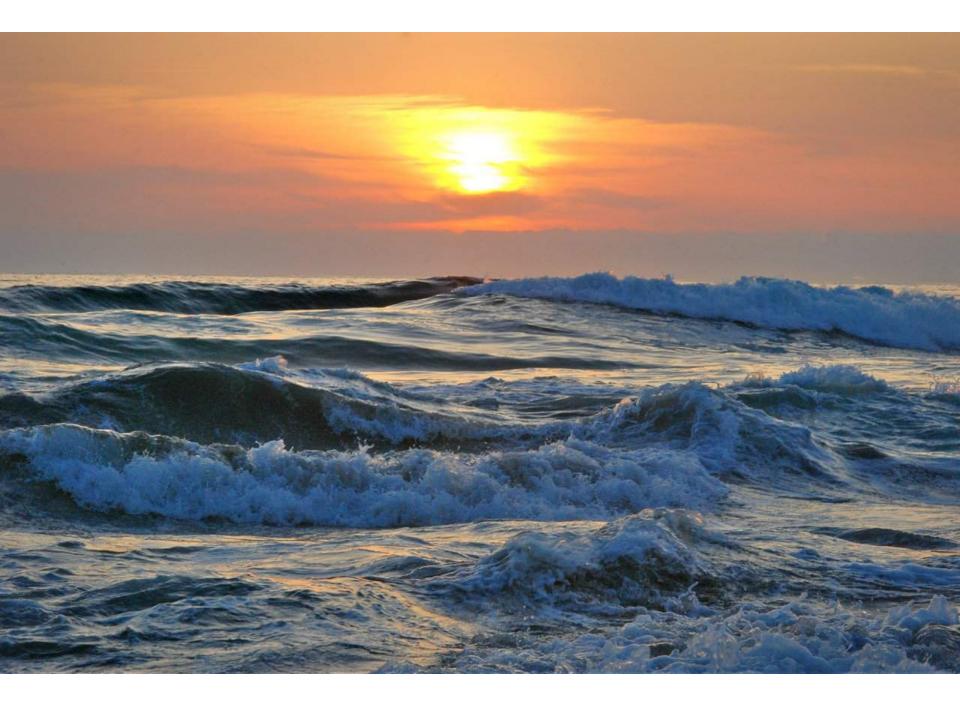
[Forget Rotation For Now]

Basic Driving Force?: Gravity, g

cf., Pendulum:







Two Basic Modes of Movements:

Gravity-Wave Speed: $c_g = (gh)^{1/2}$

Rotation Velocity : $v_R = l \Omega$ (generalized) for a spatial distance, *l* (length of arm)

Two Basic Modes of Movements: Rotation Velocity Velocity Cg **Gravity Wave** spatial distance, l

Two Basic Modes of Movements: Rotation Velocity Velocity C_{g} **Gravity Wave** spatial distance, l transition scale

Gravity Wave = Rotation Velocity

$$c_g = v_R$$
$$c_g = l \Omega$$

transition scale: $l = \Omega/c_g$: typical scale of the weather

relative transition scale (to radius):

$$l/r = \Omega/rc_g$$

Problem with estimating the Transition Scale:

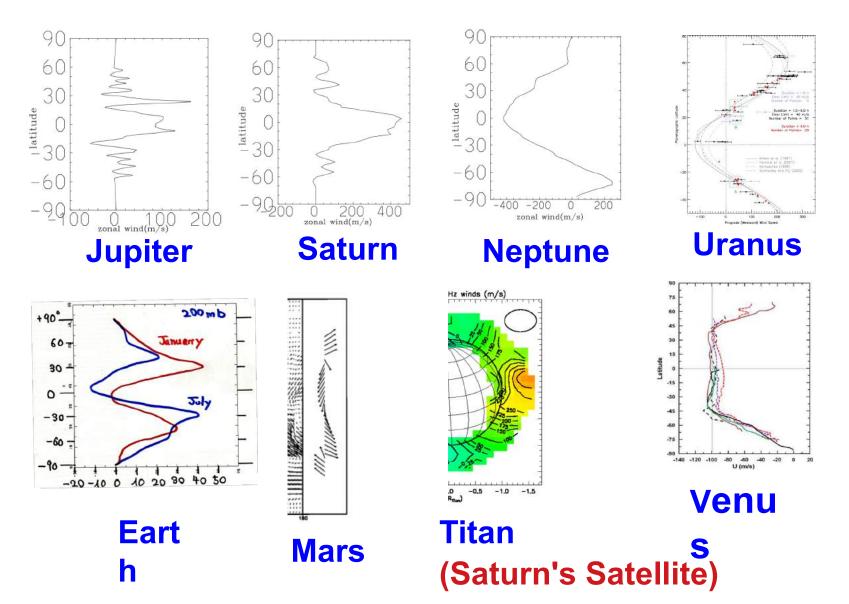
h = ? or $c_g = ?$ NB: h is only an *effective* depth of the atmosphere

Set : c_g = 50m/s: typical value in the Earth's atmosphere

Nondimensional Parameters:

Planet	Rossby number	Transition Scale (relative to the radius)
Venus	-55	-27
Earth	0.02	0.1
Mars	0.04	0.2
Jupiter	0.008	0.0041
Saturn	0.02	0.0048
Uranus	0.04	0.020
Neptune	0.08	0.021
Titan	17	4.3

Eastward Winds with Latitudes: In Increasing Order of the Transition Scale:



Conclusions (1):

Weather of the atmosphere is characterized by the scale in which the gravity-wave speed, $c_q = (gh)^{1/2}$, becomes equal to the rotational velocity, $V_{R} = l\Omega$.

i.e., **Transition Scale:** $l = \Omega/c_{a}$



Conclusions (2):

As a result, as the planet rotates faster (slower), the scale that characterizes the weather becomes smaller (larger)

Conclusions (3):

With slower rotations: weather is smoother with less dependence in longitude (e.g., Hadley cell)

<u>With a very slow rotation:</u> the winds blow very fast eastwards (in the direction of rotation)

Conclusions (4):

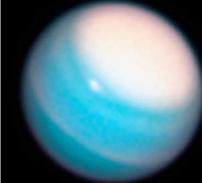
- With faster rotations:
- the weather realizes in smaller
 - scales and evolves faster
- number of jet streams (eastwardwind peaks) increases
- more features are found in the longitudinal direction

Planets with Atmospheres In Increasing Order of the Rossby number:









Jupiter



Earth

Uranus





Mars



Titan (Saturn's Satellite)



Venus

Wisdom of This Talk:

"Simplify, Simplify Simplify"

(Henry David Thoreau, *Walden*)